

$a_1(1260)$

$I^G(J^{PC}) = 1^-(1^{++})$

THE $a_1(1260)$

Written March 1998 by S. Eidelman (Novosibirsk).

The main experimental data on the $a_1(1260)$ may be grouped into two classes:

(1) **Hadronic Production.** This comprises diffractive production with incident π^- (DAUM 80, 81B) and charge-exchange production with low-energy π^- (DANKOWYCH 81, ANDO 92). The 1980's experiments explain the $I^G LJ^P = 1^+ S0^+$ data using a phenomenological amplitude consisting of a rescattered Deck amplitude plus a direct resonance-production term. They agree on an $a_1(1260)$ mass of about 1270 MeV and a width of 300–380 MeV. ANDO 92 finds rather lower values for the mass (1121 MeV) and width (239 MeV) in a partial-wave analysis based on the isobar model of the $\pi^+\pi^-\pi^0$ system. However, in this analysis, only Breit-Wigner terms were considered.

(2) **τ decay.** Five experiments reported good data on $\tau \rightarrow a_1(1260)\nu_\tau \rightarrow \rho\pi\nu_\tau$ (RUCKSTUHL 86, SCHMIDKE 86, ALBRECHT 86B, BAND 87, and ACKERSTAFF 97R). They are somewhat inconsistent concerning the $a_1(1260)$ mass, which can, however, be attributed to model-dependent systematic uncertainties (BOWLER 86, ALBRECHT 93C, ACKERSTAFF 97R). They all find a width greater than 400 MeV.

The discrepancies between the hadronic- and τ -decay results have stimulated several reanalyses. BASDEVANT 77, 78 used the early diffractive dissociation and τ decay data and showed that they could be well reproduced with an a_1 resonance mass of 1180 ± 50 MeV and width of 400 ± 50 MeV. Later, BOWLER 86, TORNQVIST 87, ISGUR 89, and IVANOV 91

have studied the process $\tau \rightarrow 3\pi\nu_\tau$. Despite quite different approaches, they all found a good overall description of the τ -decay data with an $a_1(1260)$ mass near 1230 MeV, consistent with the hadronic data. However, their widths remain significantly larger (400–600 MeV) than those extracted from diffractive-hadronic data. This is also the case with the later OPAL experiment (ACKERSTAFF 97R). In the high statistics analysis of ACKERSTAFF 97R the models of ISGUR 89 and KUHN 90 are used to fit distributions of the 3π invariant mass as well as the 2π invariant mass projections of the Dalitz plot and neither model is found to provide a completely satisfactory description of the data. Another recent high statistics analysis of ABREU 98G obtains good description of the $\tau \rightarrow 3\pi$ data using the model of FEINDT 90 which includes the a'_1 meson, a radial excitation of the $a_1(1260)$ meson, with a mass of 1700 MeV and a width of 300 MeV.

BOWLER 88 showed that good fits to both the hadronic and the τ -decay data could be obtained with a width of about 400 MeV. However, applying the same type of analysis to the ANDO 92 data, the low mass and narrow width they obtained with the Breit-Wigner PWA do not change appreciably.

CONDO 93 found no evidence for charge-exchange photo-production of the $a_1(1260)$ (but found a clear signal of $a_2(1320)$ photoproduction). They show that it is consistent with either an extremely large $a_1(1260)$ hadronic width or with a small radiative width to $\pi\gamma$, which could be accommodated if the a_1 mass is somewhat below 1260 MeV.

a₁(1260) MASS

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
1230±40 OUR ESTIMATE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1262± 9± 7	1,2 ACKERSTAFF 97R OPAL			$E_{\text{cm}}^{\text{ee}} = 88-94,$ $\tau \rightarrow 3\pi\nu$
1210± 7± 2	2,3 ACKERSTAFF 97R OPAL			$E_{\text{cm}}^{\text{ee}} = 88-94,$ $\tau \rightarrow 3\pi\nu$
1211± 7	ALBRECHT 93C ARG			$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
1121± 8	4 ANDO 92 SPEC			$8\pi^- p \rightarrow \pi^+\pi^-\pi^0 n$
1242±37	5 IVANOV 91 RVUE			$\tau \rightarrow \pi^+\pi^+\pi^-\nu$
1260±14	6 IVANOV 91 RVUE			$\tau \rightarrow \pi^+\pi^+\pi^-\nu$
1250± 9	7 IVANOV 91 RVUE			$\tau \rightarrow \pi^+\pi^+\pi^-\nu$
1208±15	ARMSTRONG 90 OMEG 0			$300.0 pp \rightarrow pp\pi^+\pi^-\pi^0$
1220±15	8 ISGUR 89 RVUE			$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
1260±25	9 BOWLER 88 RVUE			
1166±18±11	BAND 87 MAC			$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
1164±41±23	BAND 87 MAC			$\tau^+ \rightarrow \pi^+\pi^0\pi^0\nu$
1250±40	8 TORNQVIST 87 RVUE			
1046±11	ALBRECHT 86B ARG			$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
1056±20±15	RUCKSTUHL 86 DLCO			$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
1194±14±10	SCHMIDKE 86 MRK2			$\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$
1240±80	10 DANKOWY... 81 SPEC 0			$8.45\pi^- p \rightarrow n3\pi$
1280±30	10 DAUM 81B CNTR			$63,94\pi^- p \rightarrow p3\pi$
1041±13	11 GAVILLET 77 HBC +			$4.2 K^- p \rightarrow \Sigma 3\pi$

¹ Uses the model of KUHN 90.² Supersedes AKERS 95P.³ Uses the model of ISGUR 89.⁴ Average and spread of values using 2 variants of the model of BOWLER 75.⁵ Reanalysis of RUCKSTUHL 86.⁶ Reanalysis of SCHMIDKE 86.⁷ Reanalysis of ALBRECHT 86B.⁸ From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.⁹ From a combined reanalysis of ALBRECHT 86B and DAUM 81B.¹⁰ Uses the model of BOWLER 75.¹¹ Produced in K^- backward scattering.

$a_1(1260)$ WIDTH

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
250 to 600 OUR ESTIMATE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
621 ± 32 ± 58	12,13 ACKERSTAFF 97R OPAL			$E_{\text{cm}}^{\text{ee}} = 88-94,$ $\tau \rightarrow 3\pi\nu$
457 ± 15 ± 17	13,14 ACKERSTAFF 97R OPAL			$E_{\text{cm}}^{\text{ee}} = 88-94,$ $\tau \rightarrow 3\pi\nu$
446 ± 21	ALBRECHT 93C ARG			$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
239 ± 11	ANDO 92 SPEC			$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
266 ± 13 ± 4	15 ANDO 92 SPEC			$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
465 $^{+228}_{-143}$	16 IVANOV 91 RVUE			$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
298 $^{+40}_{-34}$	17 IVANOV 91 RVUE			$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
488 ± 32	18 IVANOV 91 RVUE			$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
430 ± 50	ARMSTRONG 90 OMEG 0			$300.0 pp \rightarrow pp\pi^+\pi^-\pi^0$
420 ± 40	19 ISGUR 89 RVUE			$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
396 ± 43	20 BOWLER 88 RVUE			
405 ± 75 ± 25	BAND 87 MAC			$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
419 ± 108 ± 57	BAND 87 MAC			$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
521 ± 27	ALBRECHT 86B ARG			$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
476 $^{+132}_{-120} \pm 54$	RUCKSTUHL 86 DLCO			$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
462 ± 56 ± 30	SCHMIDKE 86 MRK2			$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
380 ± 100	21 DANKOWY... 81 SPEC 0			$8.45 \pi^- p \rightarrow n3\pi$
300 ± 50	21 DAUM 81B CNTR			$63.94 \pi^- p \rightarrow p3\pi$
230 ± 50	22 GAVILLET 77 HBC +			$4.2 K^- p \rightarrow \Sigma 3\pi$

12 Uses the model of KUHN 90.

13 Supersedes AKERS 95P

14 Uses the model of ISGUR 89.

15 Average and spread of values using 2 variants of the model of BOWLER 75.

16 Reanalysis of RUCKSTUHL 86.

17 Reanalysis of SCHMIDKE 86.

18 Reanalysis of ALBRECHT 86B.

19 From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.

20 From a combined reanalysis of ALBRECHT 86B and DAUM 81B.

21 Uses the model of BOWLER 75.

22 Produced in K^- backward scattering.

$a_1(1260)$ DECAY MODES

Mode	Fraction (Γ_i/Γ)
$\Gamma_1 \rho\pi$ [D/S amplitude ratio = -0.100 ± 0.028]	dominant
$\Gamma_2 \pi\gamma$	seen
$\Gamma_3 \pi(\pi\pi)_S$ -wave	possibly seen

 $a_1(1260)$ PARTIAL WIDTHS

$\Gamma(\pi\gamma)$	Γ_2
<u>VALUE (keV)</u> 640±246	<u>DOCUMENT ID</u> ZIELINSKI <u>TECN</u> 84C SPEC <u>COMMENT</u> $200 \pi^+ Z \rightarrow Z3\pi$

 D -wave/ S -wave AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
-0.10 ± 0.02 ± 0.02	23,24 ACKERSTAFF 97R OPAL	E _{cm} ^{ee} = 88–94, $\tau \rightarrow 3\pi\nu$	

23 Uses the model of ISGUR 89.

24 Supersedes AKERS 95P

 $a_1(1260)$ BRANCHING RATIOS

$\Gamma(\pi(\pi\pi)_S$ -wave)/ $\Gamma(\rho\pi)$	Γ_3/Γ_1
<u>VALUE</u> 0.003±0.003	<u>DOCUMENT ID</u> 25 LONGACRE 82 RVUE <u>TECN</u>

• • • We do not use the following data for averages, fits, limits, etc. • • •

25 Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from GAVILLET 77, DAUM 80, and DANKOWYCH 81.

 $a_1(1260)$ REFERENCES

ACKERSTAFF 97R ZPHY C75 593	K. Ackerstaff+	(OPAL Collab.)
AKERS 95P ZPHY C67 45	+Alexander, Allison, Ametewee+	(OPAL Collab.)
ALBRECHT 93C ZPHY C58 61	+Ehrlichmann, Hamacher+	(ARGUS Collab.)
ANDO 92 PL B291 496	+Imai+	(KEK, KYOT, NIRS, SAGA, INUS, AKIT)
IVANOV 91 ZPHY C49 563	+Osipov, Volkov	(JINR)
ARMSTRONG 90 ZPHY C48 213	+Benayoun, Beusch	(WA76 Collab.)
KUHN 90 ZPHY C48 445	J.H. Kuhn, Santamaría+	(MPIM)
ISGUR 89 PR D39 1357	+Morningstar, Reader	(TNT)
BOWLER 88 PL B209 99	+Camporesi, Chadwick, Delfino+	(OXF)
BAND 87 PL B198 297	+Donker, Gabriel, Edwards+	(MAC Collab.)
TORNQVIST 87 ZPHY C36 695	+Stroynowski, Atwood, Barish+	(HELS)
ALBRECHT 86B ZPHY C33 7	+Abrams, Matteuzzi, Amidei+	(ARGUS Collab.)
RUCKSTUHL 86 PRL 56 2132	+Berg, Chandee, Cihangir+	(DELCO Collab.)
SCHMIDKE 86 PRL 57 527	Dankowych+	(Mark II Collab.)
ZIELINSKI 84C PRL 52 1195	+Hertzberger+	(ROCH, MINN, FNAL)
LONGACRE 82 PR D26 83	+Hertzberger+	(BNL)
DANKOWY... 81 PRL 46 580	+Blockzijl, Engelen+	(TNTO, BNL, CARL, MCGI, OHIO)
DAUM 81B NP B182 269	+Game, Aitchison, Dainton	(AMST, CERN, CRAC, MPIM, OXF+)
DAUM 80 PL 89B 281		(AMST, CERN, CRAC, MPIM, OXF+) JP
GAVILLET 77 PL 69B 119		(AMST, CERN, NIJM, OXF) JP
BOWLER 75 NP B97 227		(OXFTP, DARE)

OTHER RELATED PAPERS

ABREU	98G	PL B (to be publ.)	P. Abreu+	(DELPHI Collab.)
CERN-EP/98-14				
BOLONKIN	95	PAN 58 1535 Translated from YAF 58 1628.	+Vladimirskii, Erofeeva+	(ITEP)
WINGATE	95	PRL 74 4596	+De Grand	(COLO, FSU)
CONDO	93	PR D48 3045	+Handler, Bugg+	(SLAC Hybrid Collab.)
FEINDT	90	ZPHY C48 681	M. Feindt	(HAMB)
IIZUKA	89	PR D39 3357	+Koibuchi, Masuda	(NAGO, IBAR, TSUK)
TORNQVIST	87	ZPHY C36 695		(HELS)
BOWLER	86	PL B182 400		(OXF)
BASDEVENT	78	PRL 40 994	+Berger	(FNAL, ANL) JP
BASDEVENT	77	PR D16 657	+Berger	(FNAL, ANL) JP
ADERHOLZ	64	PL 10 226	+ (AACH3, BERL, BIRM, BONN, DESY, HAMB+)	
GOLDHABER	64	PRL 12 336	+Brown, Kadyk, Shen+	(LRL, UCB)
LANDER	64	PRL 13 346A	+Abolins, Carmony, Hendricks, Xuong+	(UCSD) JP
BELLINI	63	NC 29 896	+Fiorini, Herz, Negri, Ratti	(MILA)
